

Satellite Constellation Orbit Design to Enable a Space-Based Radio Interferometer

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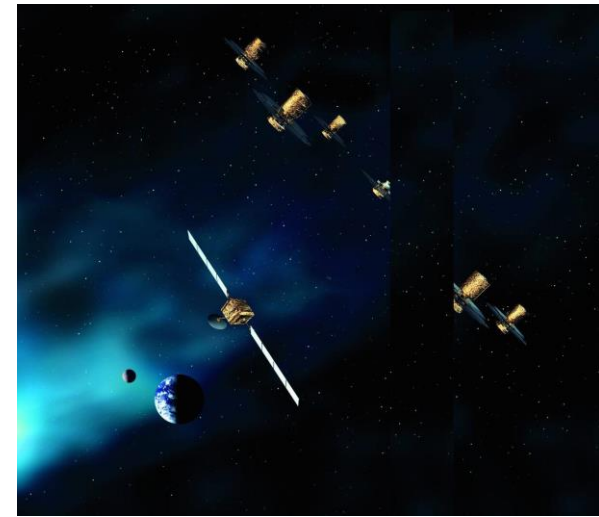
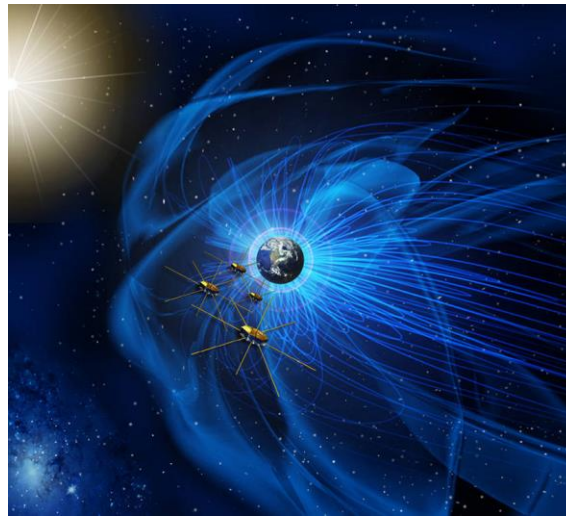
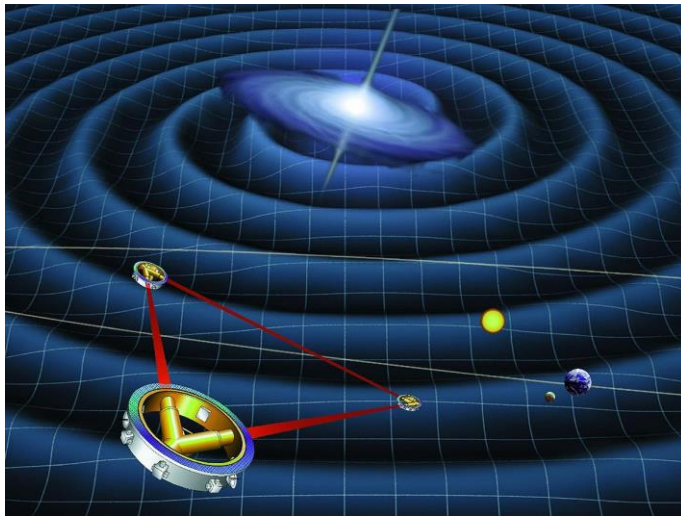
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Introduction



- **Networked constellation** missions of small spacecraft benefit from lower cost, increased robustness, and enable novel types of missions
- Advancements in **autonomy** make operations of constellations more manageable and less costly
- **Space-based radio interferometers**, where a number of satellites in a (networked) constellation act as an array of radio telescopes can achieve resolutions impossible to achieve with Earth-based interferometers
- Mothership in a reference orbit can act as a **relay** for the constellation
- What **technology** exists today and what technology is still needed to enable these types of missions?

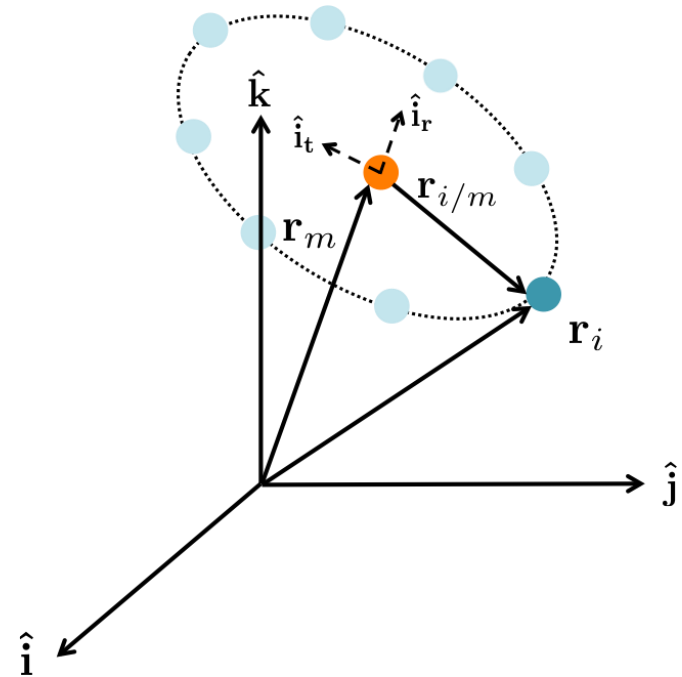
Constellation Design Methods

- Mothership (reference) spacecraft
- N daughter spacecraft
- Relative to the mothership, in a RTN frame

$$\mathbf{r}_{i/m} = \begin{bmatrix} x_i \hat{\mathbf{i}}_r \\ y_i \hat{\mathbf{i}}_t \\ z_i \hat{\mathbf{i}}_n \end{bmatrix}$$

$$\mathbf{v}_{i/m} = \begin{bmatrix} \dot{x}_i \hat{\mathbf{i}}_r \\ \dot{y}_i \hat{\mathbf{i}}_t \\ \dot{z}_i \hat{\mathbf{i}}_n \end{bmatrix}$$

- Any formation design strategy must begin with a **method to predict and analyze** both the **absolute** as well as **relative** motion of spacecraft.



- **Goal:** Design simple and accurate design strategies that can quickly assess the best geometry design for N spacecraft constellation

1. Constellation design using Linear Dynamics (CLD)

- Analytical framework, especially useful for initial design stages

2. Constellation design using Invariant Manifold (CIM)

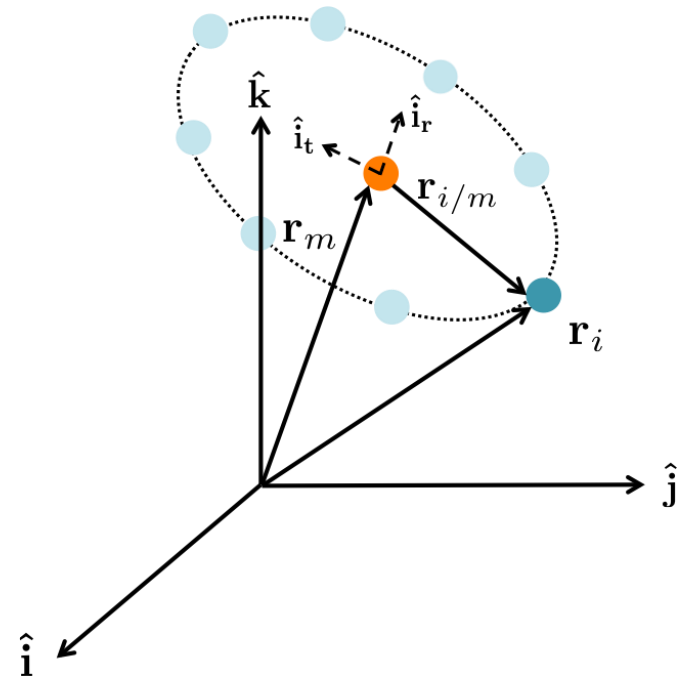
- Advantageous when considering higher order dynamics

Constellation design using Linear Dynamics (CLD)

- Clohessy-Wiltshire (CW)**

- Assumes mothership is in a circular orbit about the central body
- Write relative motion equations assuming two-body motion
- Use a binomial expansion to first order, and assume $r_{i/m}$ is small

$$\begin{aligned}\ddot{x}_i &= 2\omega\dot{y}_i + 3\omega^2 x_i \\ \ddot{y}_i &= -2\omega\dot{x}_i \\ \ddot{z}_i &= -\omega^2 z_i\end{aligned}$$



- Analytical Solution**

$$\begin{aligned}x_i(t) &= 2(2x_{i_0} + \dot{y}_{i_0}/\omega) - (3x_{i_0} + 2\dot{y}_{i_0}/\omega) \cos \xi + (\dot{x}_{i_0}/\omega) \sin \xi \\ y_i(t) &= (y_{i_0} - 2\dot{x}_{i_0}/\omega) - 3(2x_{i_0} + \dot{y}_{i_0}/\omega)\xi + (2\dot{x}_{i_0}/\omega) \cos \xi + 2(3x_{i_0} + 2\dot{y}_{i_0}/\omega) \sin \xi \quad (7) \\ z_i(t) &= z_{i_0} \cos \xi + (\dot{z}_{i_0}/\omega) \sin \xi\end{aligned}$$

$$\xi = \omega t$$

Constellation design using Linear Dynamics (CLD)

- **Clohessy-Wiltshire (CW)**

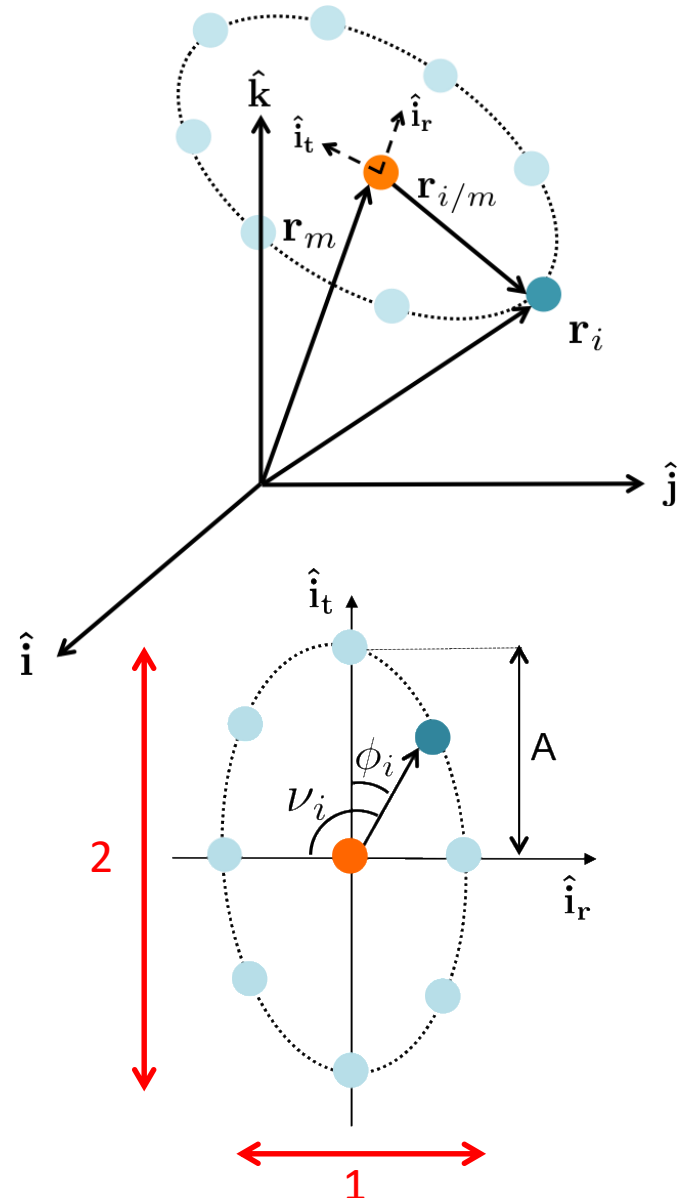
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- **Periodic Motion**

$$\begin{aligned}x_i(t) &= \frac{1}{2}A \sin(\alpha + \phi_i(t)) \\ y_i(t) &= A \cos(\alpha + \phi_i(t)) + y_c \\ z_i(t) &= B \sin(\beta + \phi_i(t))\end{aligned}$$

In-plane motion is given by **2x1** ellipse or **ring**, where size is a function of the **eccentricity**



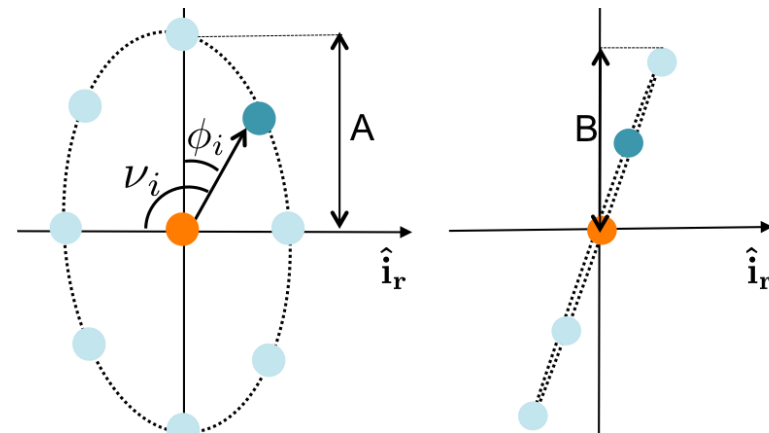
Using CLD for Constellation Design

- Relative motion for ***N spacecraft***:

For each ring in a constellation:

For each spacecraft on a ring:

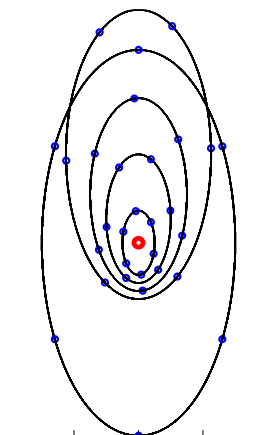
$$\begin{aligned} x_{i_j}(t) &= \frac{1}{2} A_i \sin(\phi_{i_j}(t)) \\ y_{i_j}(t) &= A_i \cos(\phi_{i_j}(t)) + y_{c_i} \\ z_{i_j}(t) &= B_i \sin(\beta_i + \phi_{i_j}(t)) \end{aligned}$$



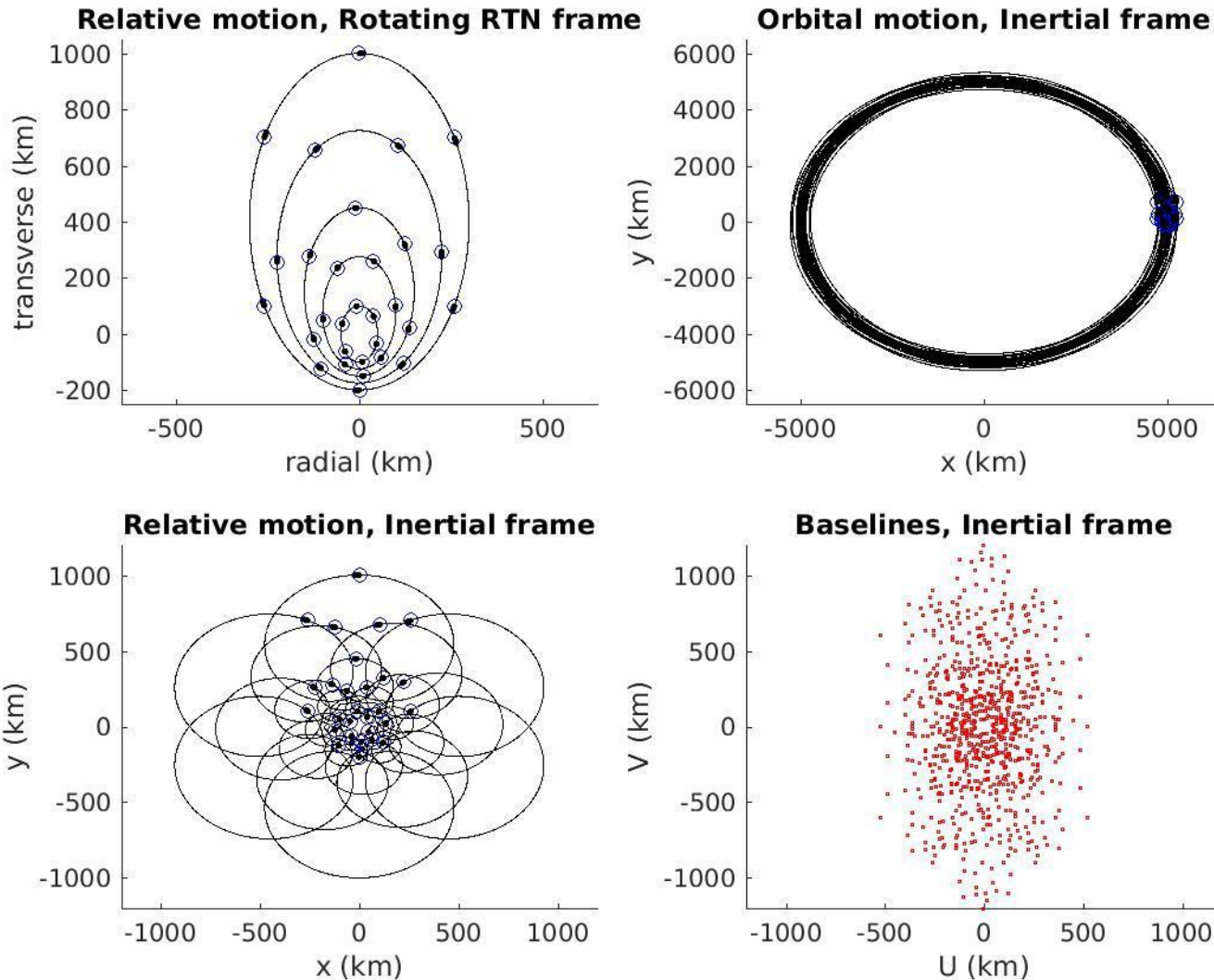
- 5N parameters**: **4N geometry** parameters and **N phasing** parameters

Formation Parameters	Ring Parameters	
	In-plane	Out-of-plane
ω = orbit period	A = ring size	B = max. displacement
n_r = n° rings	ϕ = angular displacement	β = ring orientation
$n_{sc/r}$ = n° of spacecraft/ring	y_c = center of ring	

- Design **rings** of varying sizes and centers, mimicking a gear-like movement which allows for optimum science target (baseline) coverage

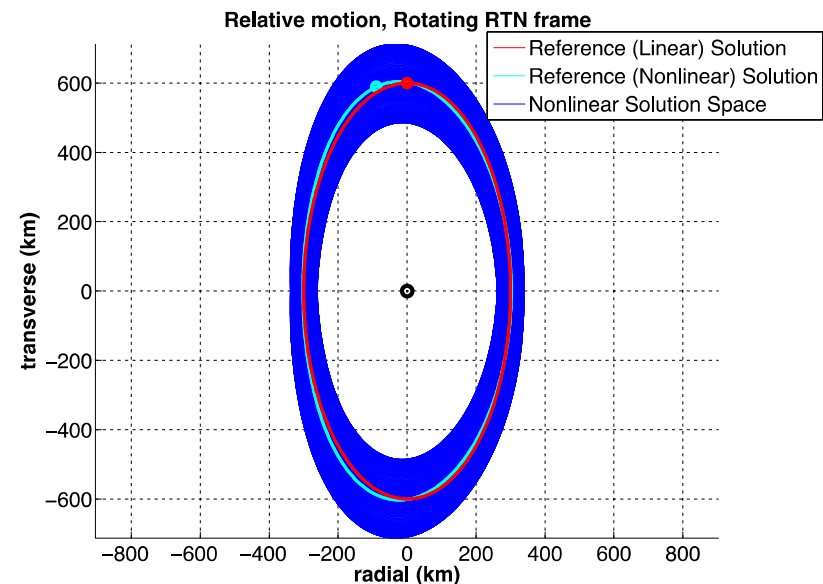
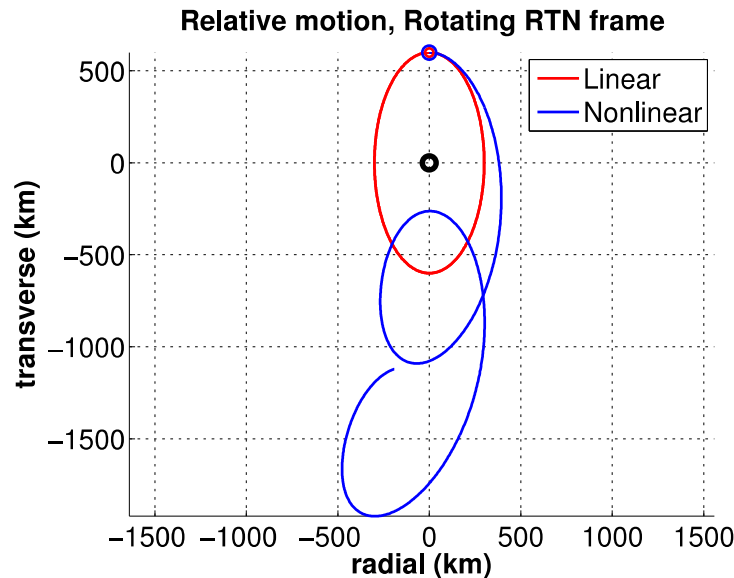


CLD: Example Constellation around Moon



Converting from Linear to Nonlinear Motion

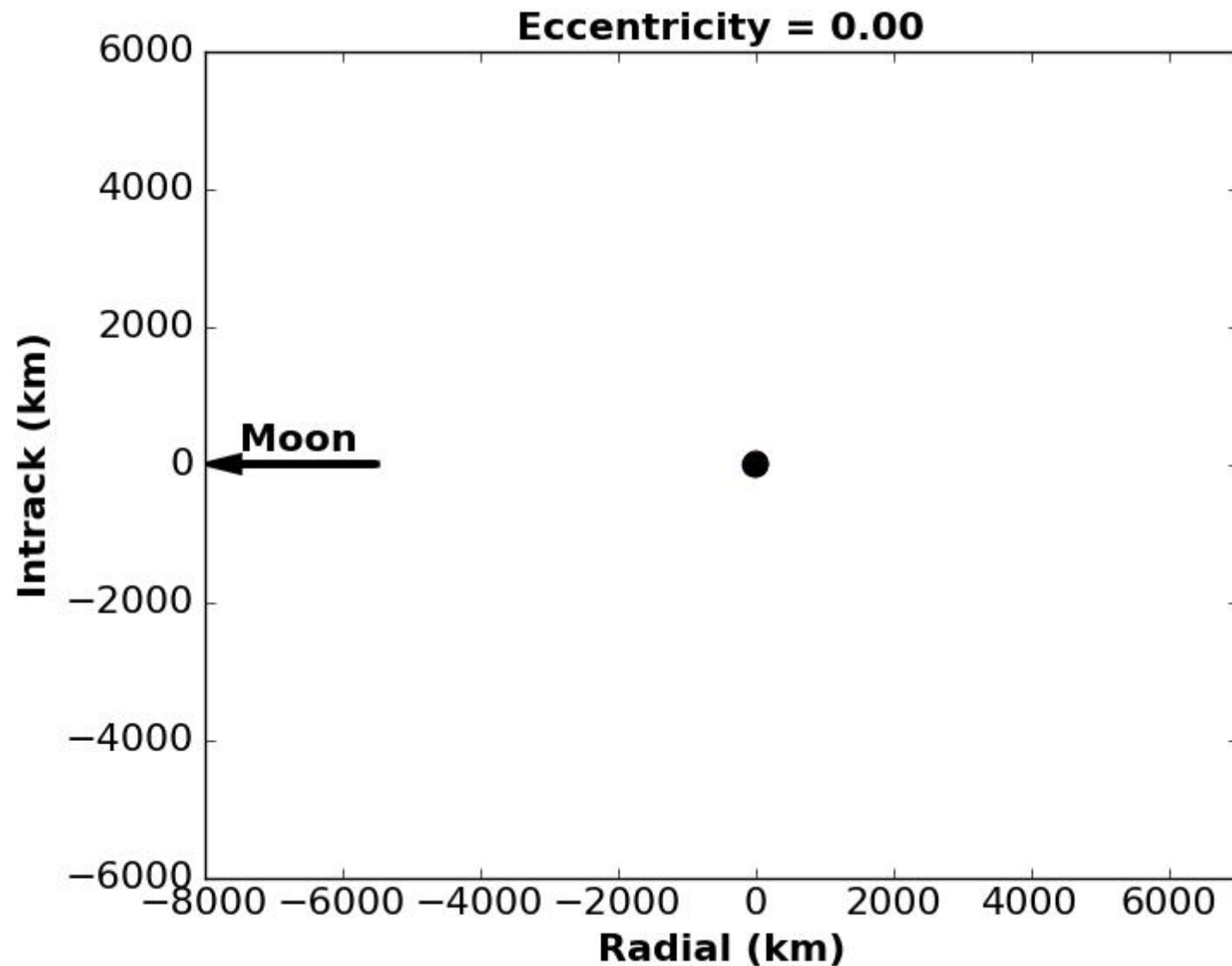
- Fix ***semi-major axis*** to the same one as the mothership
- Depending on phase at which the conversion is performed, greater discrepancies from the linear solution might appear.



- ***Minimum deviation*** between linear and nonlinear occurs at $\varphi=0^\circ$ ($v=90^\circ$)
- Make conversion at $\varphi=0^\circ$ to obtain ring geometry, then can distribute spacecraft at any desired phase

Linear model can be used to initially design the constellation, because when converting to the **nonlinear model** we are able to reproduce the **same design**

Eccentricity Variation in Nonlinear Model



2. Constellation design using Invariant Manifolds (CIM)

- Use **dynamical system theory** to **excite center eigenvectors** of the reference path in different directions to create the constellation
- Motion of daughter spacecraft can be **linearized** about a reference path

$$\dot{\mathbf{x}}_i = \mathbf{A}(t)\mathbf{x}_i \quad \mathbf{A}(t) = \begin{bmatrix} \mathbf{0}_{3 \times 3} & \mathbf{I}_{3 \times 3} \\ \frac{\partial^2 U}{\partial \mathbf{r}_m^2} & \mathbf{0}_{3 \times 3} \end{bmatrix}$$

- Variations in initial state to final state are given by STM

$$\mathbf{x}_i(t) = \Phi(t, t_0)\mathbf{x}_i(t_0) \quad \dot{\Phi}(t, t_0) = \mathbf{A}(t)\Phi(t, t_0)$$

- Monodromy matrix: STM propagated for one period
 - **Eigenvalues** λ_i , $i = 1, \dots, 6$ give stability of periodic (reference) orbit
 - **Eigenvector** \mathbf{e}_i , $i = 1, \dots, 6$ are used to excite relative motion in specific directions
- Relative motion can be generated via

$$\mathbf{x}_i(\tau, \boldsymbol{\theta}_i, \boldsymbol{\varepsilon}_i) = \sum_{k=1}^n \varepsilon_{i_k} \left(\cos(\theta_{i_k}) \text{Re}[\hat{\mathbf{e}}_k(\tau)] - \sin(\theta_{i_k}) \text{Im}[\hat{\mathbf{e}}_k(\tau)] \right)$$

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- For **two-body motion** $|\lambda_i|=1$
 - 2 form a complex conjugate pair
 - 2 are repeated strictly real vectors (monodromy matrix degenerate)
 - 2 eigenvectors are unique and strictly real

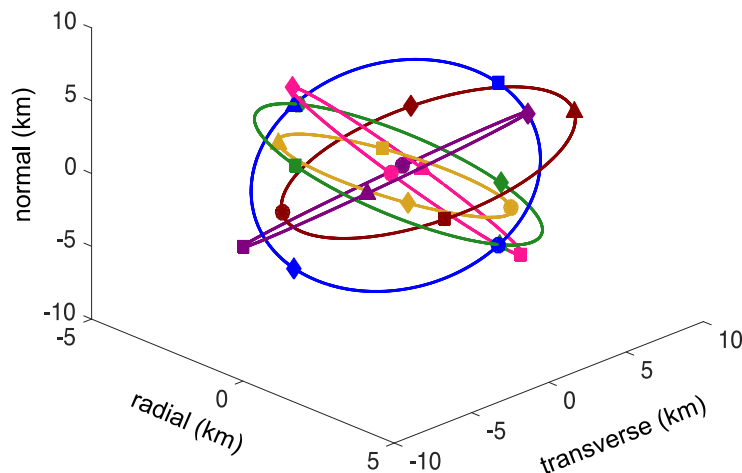
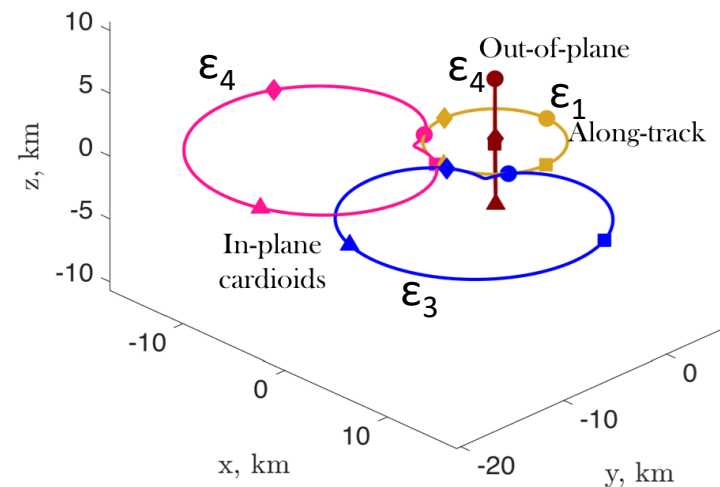
4 geometry parameters and **1 phasing** parameter

$$x_i(\varepsilon_i, \theta_i) = \varepsilon_{i_1} \hat{\mathbf{e}}_1 + \varepsilon_{i_2} \hat{\mathbf{e}}_2 + \varepsilon_{i_3} \hat{\mathbf{e}}_3 + \varepsilon_{i_4} \left(\cos(\theta_{i_4}) \text{Re}[\hat{\mathbf{e}}_4] - \sin(\theta_{i_4}) \text{Im}[\hat{\mathbf{e}}_4] \right)$$

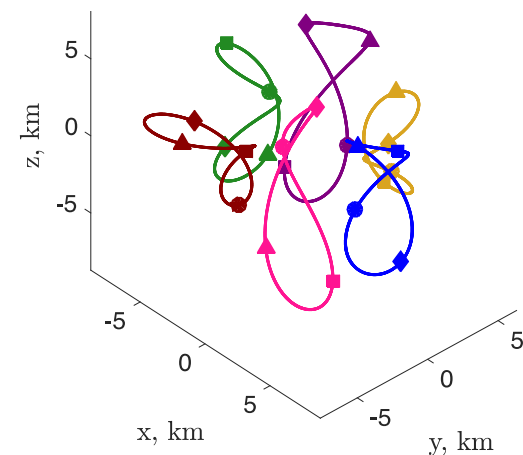
Example Constellation in GEO Graveyard using CIM

- 6 s/c constellation with the purpose of observing coronal mass ejections from the Sun
- GEO Graveyard 25 hour period

i	ϵ_{i1} (km)	ϵ_{i2} (km)	ϵ_{i3} (km)	ϵ_{i4} (km)	θ_{i4} (deg)
1	3	0	-2	2.5	150
2	0	2	-2	2	240
3	-2.5	2	0	3	60
4	-2	0	2.5	3	180
5	0	-2	2	2	300
6	3	-2.5	0	3	120



Relative frame fixed at reference (RTN)



Relative frame inertial equatorial axis

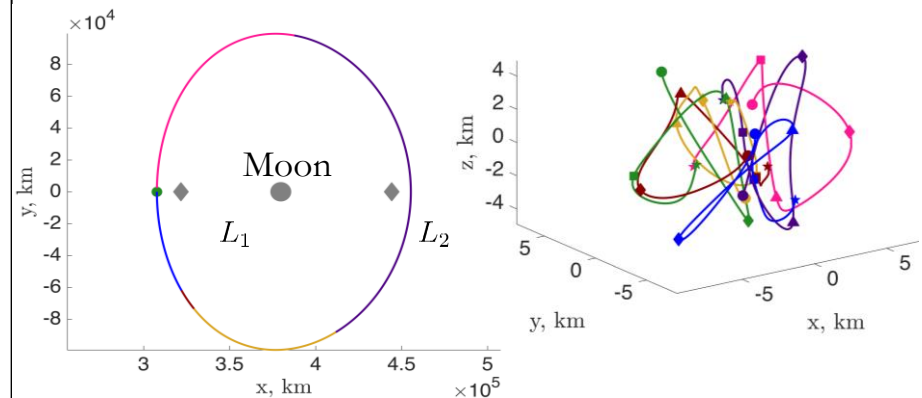
Comparison Between CLD and CIM

CLD

- **Analytical** solution exists in linear model, which is advantageous for initial design stages
- For small eccentricity, conversion to two-body model can be made while preserving the same linear geometry
- For higher order dynamics, stationkeeping costs need to be taken into account
- Constellation same period
- **5N parameters**
 - 4N ring geometry: A, B, γ_c, β
 - 1N phasing: φ

CIM

- **Higher order dynamics** can be used to define the constellation
- Example: **DRO in CR3BP**

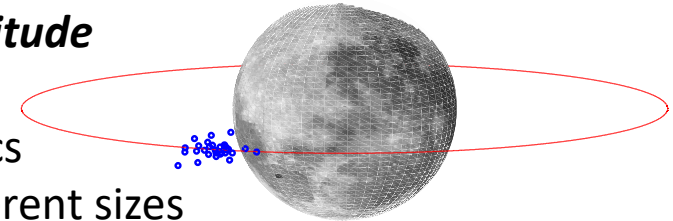


- Constellation same period
- **5N parameters**
 - 4N ring geometry: $\epsilon_1, \epsilon_2, \epsilon_3, \epsilon_4$
 - 1N phasing: θ

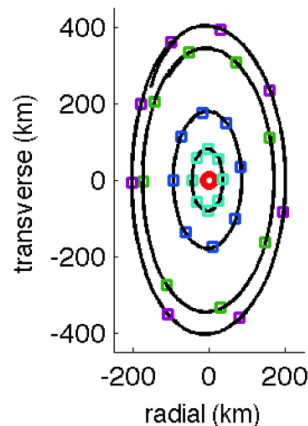
Even though both methods are derived using **different approaches**, the **same constellation design** can be achieved with either method

Lunar Orbiting Constellation Example

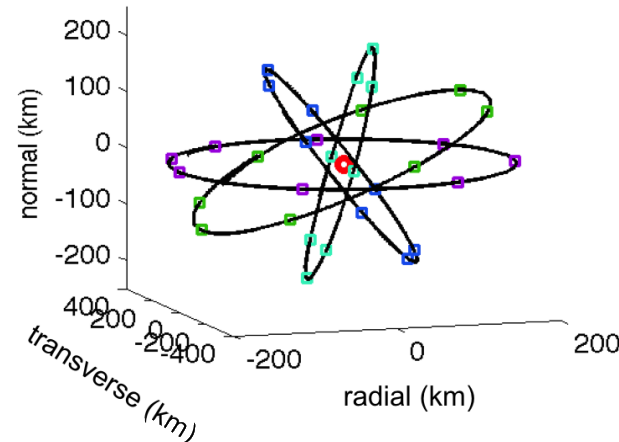
- Goal: Observation of **distant galaxies** using **radio interferometer**
- The constellation needs to be far enough from the Earth to avoid interference
- Reference orbit at circular orbit at **5,000 km lunar altitude**
- 600 km maximum baseline
- Use CLD for constellation design in two-body dynamics
 - **32 daughter s/c** distributed along **4 rings** of different sizes
 - 12kg small satellites, 1N thruster, Isp = 200 s



Motion Relative to Mothership, Planar View



Motion Relative to Mothership, 3D View



- **Operations:**

1. Deployment: Depart in ESPA ring from mothership, using 20 m/s
2. Reconfiguration: 20 m/s allowed over 6 month mission time span
3. Baseline Coverage: Observe maximum celestial sphere
4. Stationkeeping Costs

Reconfiguration

- Reconfiguration allows for more baseline acquisition, which allows for *more science*
- Need orbit **period to be constant**, so constellation does not drift apart
- Maneuvering direction **perpendicular to velocity**

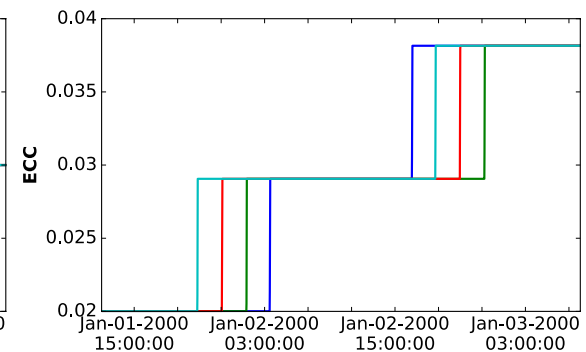
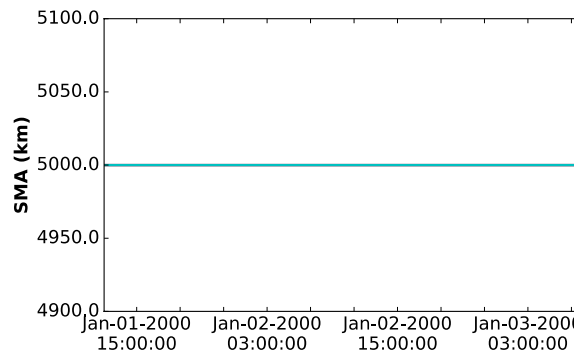
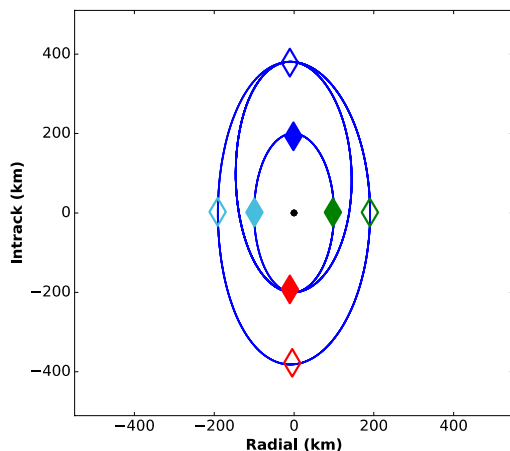
$$\ddot{\mathbf{r}}_i = -\mu \frac{\mathbf{r}_i}{r_i^3} + \mathbf{f}_i \quad E_i = \frac{1}{2}(\mathbf{v}_i^T \mathbf{v}_i) - \mu(\mathbf{r}_i^T \mathbf{r}_i)^{-1/2} \quad \dot{E}_i = \mathbf{v}_i^T \mathbf{f}_i = 0 \iff \mathbf{f}_i \perp \mathbf{v}_i$$

- In-plane and out-of-plane change

$$\mathbf{f}_i = \frac{T}{m} \left[\frac{k_1 \hat{\mathbf{v}}_i \times \hat{\mathbf{h}}_i + k_2 \hat{\mathbf{h}}_i}{\|k_1 \hat{\mathbf{v}}_i \times \hat{\mathbf{h}}_i + k_2 \hat{\mathbf{h}}_i\|} \right]$$

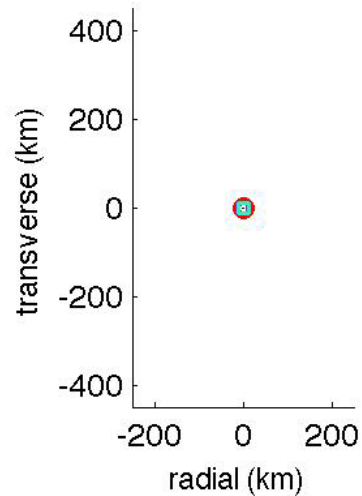
k1 in-plane thrust
k2 out-of-plane thrust

- **Example:** 4 spacecraft reconfigure to larger ring
 - Maneuver occurs at same location on relative orbit at $\mathbf{v}=\pm 90^\circ$
 - Reconfigure as many times within a ΔV budget

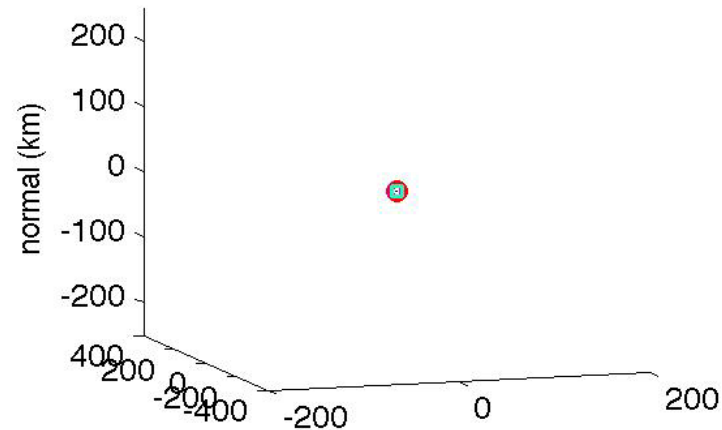


Full Mission Simulation

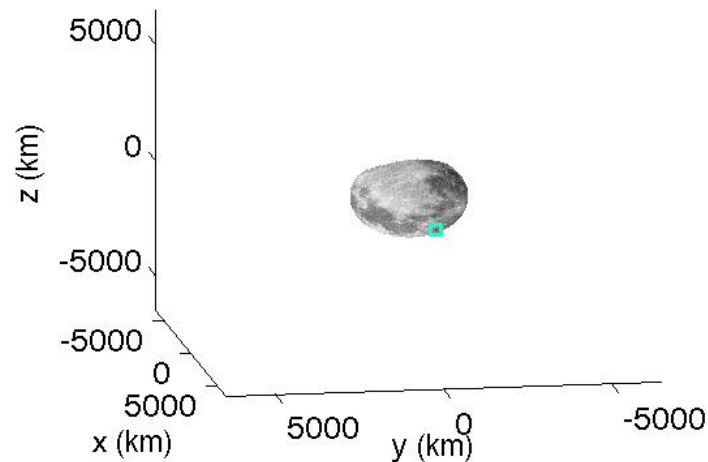
Motion Relative to Mothership, Planar View



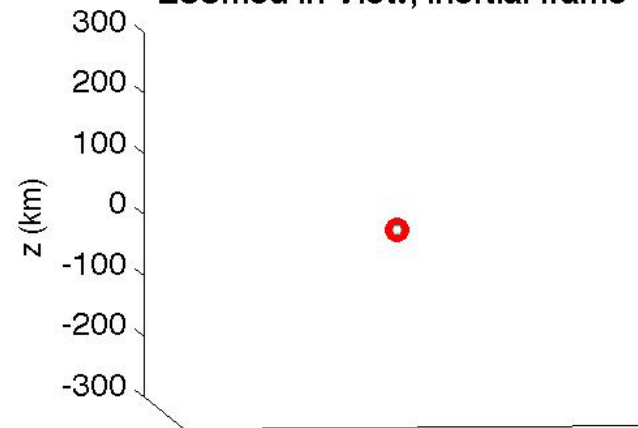
Motion Relative to Mothership, 3D View



Orbital motion, Inertial frame

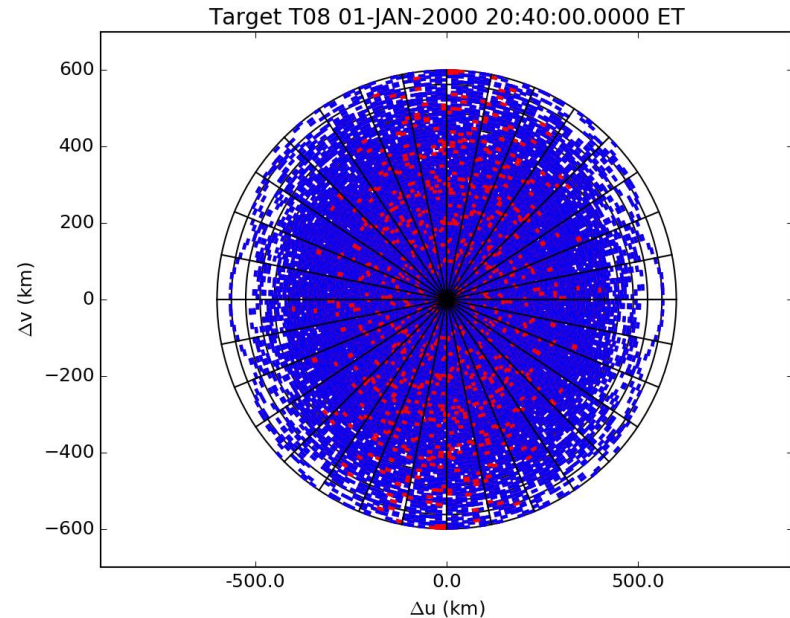
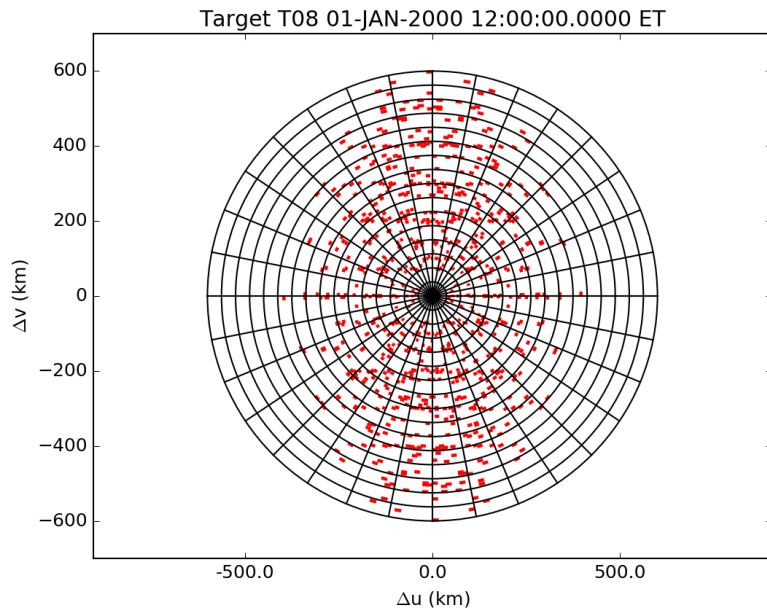


Zoomed in View, Inertial frame



Baselines Achieved in the Celestial Sphere

- The formation design is driven by **adequate coverage** of collection targets
- Coverage is **diversity of baselines** formed by individual spacecraft pairs
- Baseline is projection of the relative position vector from one spacecraft to another, into the plane perpendicular to the direction of a target
- Example target at RA 0° and DEC 45° after final reconfiguration



- For entire celestial sphere, constellation achieves 98% coverage.

Conclusion

- **Networked constellations** can enable novel types of missions, at lower costs and increased robustness
 - Spaced-based radio interferometers
- **Constellation design methods**
 1. Constellation design using Linear Dynamics (**CLD**)
 - Analytical solution, great for initial design stage
 - Can replicate linear design in nonlinear (two-body) model
 - Stationkeeping costs to account for higher-order dynamics
 2. Constellation design using Invariant Manifolds (**CIM**)
 - Requires integration of state with STM
 - Higher-order dynamics can be included in design, avoiding large stationkeeping costs
- **Operations**
 - Deployment
 - Reconfiguration to allow more science acquisition
 - Simple algorithm that can be implemented on board for autonomous ops.
 - Valid for low-thrust or high-thrust
 - Maximum baseline coverage



Jet Propulsion Laboratory